#### Chemistry 331 Introduction to Quantum Chemistry

Instructor:	Derek Leaist (dleaist@stfx.ca), Office PSC 3072, Lab PSC 3020
Lectures:	Monday 11:15 am, Tuesday 1:15 pm, Thursday 12:15 pm (PSC 3046)
Tutorials:	Wednesday 2:15 pm (AX 23A)

**General Description:** Chemistry 331 introduces the basic ideas and applications of quantum theory, quantum chemistry, and statistical mechanics, emphasizing energy levels and how they are occupied. The postulates of quantum mechanics are developed and used to solve a variety of problems, including thermal radiation, confined particles, free particles, molecular vibration, molecular rotation, and atomic orbitals, with spectroscopic and other applications. The methodology and interpretive nature of quantum mechanics are stressed and the connection between theory and experiment is emphasized.

Prerequisite: Chemistry 232

Course Notes:	pdf copies provided
Textbook:	pdf copies of textbooks available
Assignments:	About eight problem sets will be assigned to cover the course material.
Term Tests:	Wednesday 13 October, 1:15 – 5:00 pm, AX 23A Wednesday 24 November, 2:15 – 5:00 pm, AX 23A
Final Exam:	TBA (in the Christmas exam period)
Office Hours:	Mon. 9:30 to 10:30 am, Tues. 11:30 am to 12:30 pm, Fri. 1:00 to 2:00 am

#### **Course Outline**

- 1. **Brief Summary of Classical Mechanics:** Newton's Laws. Hamilton's equations of motion. Other coordinate systems. Math review.
- 2. **The Need for Quantum Mechanics:** Failures of classical mechanics. Thermal radiation. Low-temperature heat capacities. Photoelectric effect. Atomic spectra. De Broglie relation. Electron diffraction. Wave nature of matter. The uncertainty principle.
- 3. **Postulates of Quantum Mechanics:** Connection between classical and quantum mechanics. The postulates or laws of quantum mechanics. Stationary states.
- 4. **Some One-Dimensional Solutions:** The free particle problem. Particle-in-a-box problem. Beam-potential barrier problems. Quantum mechanical tunneling. Applications.
- 5. **Important Theorems:** Commutators. Self-adjoint or Hermitian operators. Quantum numbers. Examples.
- 6. **One-Dimensional Harmonic Oscillator:** Raising and lowering operators. Energy levels and wave functions. Properties of the oscillator. Transition selection rules. Zero-point energy. A model for a vibrating molecule. Vibrational spectroscopy and diatomic molecules. Other applications.
- 7. **Multi-Dimensional Problems:** Three-dimensional free particle, particle-in-a-box, and harmonic oscillator problems. Degenerate energy levels. Applications.
- 8. **Multi-Dimensional Problems with Spherical Symmetry:** Quantum numbers and general method for the solving the problems. The angular parts of atomic orbitals. *The Harmonic Oscillator and the Rigid Rotator:* Model for a rotating molecule; rotational spectroscopy of diatomic molecules; selection rules; bond lengths; vibrational-rotational spectra; electronic-vibrational-rotational spectra. *The Hydrogen Atom:* The radial and angular parts of the wave functions for the one-electron atom and ion; energies; spectra and selection rules; average values of properties; total probability and radial distribution functions; radius of the atom; probability density plots; directed atomic orbitals.

#### <u>Course Notes</u> available on the Chem 331 Moodle pages and at: https://people.stfx.ca/dleaist/Chem331/

C331 Part 1. Introduction C331 Part 2. **Thermal Radiation** C331 Part 3. Math Review C331 Part 4. **Classical Mechanics** C331 Part 5. **The Postulates of Quantum Mechanics** C331 Part 6. **Statistics** C331 Part 7. **One-Dimensional Systems** C331 Part 8. **Free Particles** C331 Part 9. **Rigid Rotors** C331 Part 10. **Harmonic Oscillators** C331 Part 11. **Multi-Dimensional Systems** C331 Part 12. The Hydrogen Atom

#### Chemistry 331 Course Schedule (2021)

Classes:	Tuesday 07 September to Tuesday 07 December		
Tutorials:	15 September (Bloc	ks HG/JH, Wednesdays 2:15 to 5:00 pm, AX 23A)	
	29 September		
	06 October		
	03 November		
	17 November		
	01 December		
Term Tests:	Wednesday 13 Octo Wednesday 24 Nove	ber, 2:15 pm to 5:00 pm, AX 23A mber, 2:15 pm to 5:00 pm, AX 23A	
Final Exam:	<b>TBA</b> (in the Christma	as final exam period)	
	<b>Course Mark</b>		
	Final Exam	45 %	
	Term Tests	40 %	
	Assignments	15 %	



Schaum's Outline of Physical Chemistry

2<sup>nd</sup> edition (\$25 Amazon.ca)

Clyde A. Metz

Chapters on quantum mechanics and spectroscopy, also thermodynamics, electrochemistry, kinetics, and transport properties.

concise summaries and worked problems.

Pdf versions of physical chemistry textbooks with chapters on quantum mechanics and spectroscopy are available <u>online</u>. Examples:

*Physical Chemistry*, 3<sup>rd</sup> Edition, Engel and Reid https://www.academia.edu/people/search?utf8=%E2%9C%93&q=quantum+chemistry+Engel+Reid

*Physical Chemistry*, 4<sup>th</sup> Edition, Bawendi, Alberty and Silbey http://www.slideshare.net/abahanaskasmui/physical-chemistry-4th-bawendi-alberty-silbey-library-pirate

*Physical Chemistry*, Moore https://ia802302.us.archive.org/18/items/physicalchemistr029701mbp/physicalchemistr029701mmb.pdf Why do we need "quantum" mechanics? What's wrong with classical Newtonian mechanics? *An important chemical example*:

#### Hydrogen Atom

electron and a proton

( ground-state diameter about 0.10 nm )



Ok. But **why doesn't the electron spiral inwards and merge with the proton**, forming a **subatomic particle** (diameter << 0.10 nm for H atoms)?

# Hydrogen Atom

electron and a proton

positive and negative charges attract, so:

electron proton 0.0002 nm? ( subatomic particle)

Why do we need "quantum" mechanics? What's wrong with classical Newtonian mechanics? *Another important chemical example*:

**Orbitals** Example: 1s for the ground state hydrogen atom

## But what are orbitals?



# **Historical Development of Quantum Mechanics**

At the end of the 1800s, many physicists believed that all principles of physics had been discovered. There were "gaps", such as the way radiation interacts with matter, but these were felt to be "just details".

#### "Classical " Physics

•Newton's Laws and classical mechanics

•Classical Thermodynamics

•Optics, electricity, magnetism

#### Key ideas of classical physics

- continuous observables
- no restriction on the energy of systems
- determinism Everything about a system's future is known by solving for r(t) from Newton's second law.
- wave nature of light

 $\vec{F} = m\vec{a}$ 



## The Nature of Light

1801 Thomas Young provided convincing experimental evidence for the wave nature of light. Diffraction and interference patterns were observed.



interference pattern observed with Young's double slit experiment



equivalent interference pattern from waves of water

1860s Maxwell developed four equations that unified the laws of electricity and magnetism. The speed of an electromagnetic wave predicted was the same as the speed of light that was experimentally measured.

#### Light was concluded to be an electromagnetic wave.

#### The Physics of Chemistry was not as Well Developed

- The existence of nuclei and electrons was known.
- Most of the periodic table had been developed (empirically).
- Catalogs of chemical reactions were available.
- The nature of molecules and chemical bonds were "gaps" in our knowledge.

By the beginning of the 1900s, several worrisome experimental observations could **not** be explained by classical physics!

The 'small gaps' turned into fundamental problems with classical physics and a radical new theory was needed to solve them. *A revolution was brewing!* 

## **Five Key Observations that 'Violated' Classical Physics**

Several key experimental observations could **not** be explained by classical physics led to the development of quantum mechanics.

#### 1. <u>Blackbody Radiation</u> - disobeyed classical electromagnetic theory

One of the most important observed phenomena (from a historical point of view) that made scientists question classical mechanics was blackbody radiation.

What is blackbody radiation (also called thermal radiation)?

All materials:

- continually absorb and give off radiation
- in everyday experience, most radiation is emitted as *infrared*

As materials are heated to higher temperature, the radiation emitted shifts toward higher frequencies.

red hot < white hot < blue hot
 (why?)</pre>

## What is blackbody (thermal) radiation?

<u>An object is a blackbody if it absorbs all incident radiation but</u> <u>reflects none, at all frequencies</u>. The radiation emitted by a blackbody is called blackbody radiation.



Experiments showed that blackbody radiation depends only on the temperature of the object (*but curiously, not on its composition!*).

An excellent approximation to a true blackbody radiator is arranged by cutting a small hole in the wall of a hollow cavity. No radiation entering the hole from the outside escapes – it is all absorbed by the walls. The radiation leaking out of the hole is blackbody radiation (also called "**cavity radiation**").



A plot of the intensity of the blackbody radiation as a function of the frequency of the radiation.

Many attempts were made to derive expressions for the blackbody spectrum measured in experiments.

Most resulted in expressions for the intensity that grew without bound (the dashed curve) with increasing frequency.



In the classical treatment, radiation emitted by a blackbody was assumed to be due to oscillations of electrons (like electrons in an antenna that give off radiation).

In classical physics, these oscillating systems could have any energy and could radiate any frequency.

## Planck's Interpretation and the Idea of Quantization

In 1900, Max Planck proposed a revolutionary idea:

the energy of the oscillators and the oscillation frequencies could only be <u>integer multiples</u> of certain values

 $E = n h \nu$  n = 0, 1, 2, 3...

**h** a constant of proportionality

v radiation frequency



In other words, Planck proposed that the energy of the oscillators was *quantized* and that only certain quantities of light energy could be emitted:  $E_1 = hv_1$ ,  $E_2 = hv_2$ ,  $E_3 = hv_3$ , ...

But in classical physics, physical quantities are allowed to take on a *continuous* values. "Quantization" was not initially accepted.

By assuming energy quantization, Planck derived expressions that accurately reproduced the measured spectrum of blackbody radiation using the constant  $h = 6.626 \times 10^{-34} \text{ J s}$ 

$$\rho(T,v) = \frac{8\pi h v^3}{c^3} \frac{1}{e^{hv/kT} - 1}$$





Nobel Prize for the discovery of energy quanta.

 $\nu/s^{-1}$ 

# What's going on?

classical (*incorrect*!) spectrum of thermal radiation

$$\rho(T,\nu)_{\text{classical}} = \frac{8\pi h\nu^2 kT}{c^3}$$

Planck's (*correct*!) spectrum of thermal radiation

$$\rho(T,v) = \frac{8\pi h v^3}{c^3} \frac{1}{e^{hv/kT} - 1}$$

Notice: 
$$\rho(T, \nu) = \frac{h\nu}{kT} \frac{1}{e^{h\nu/kT} - 1} \rho_{\text{classcial}}(T, \nu)$$

# What's going on?

$$\rho(T,\nu) = \frac{h\nu}{kT} \frac{1}{e^{h\nu/kT} - 1} \rho_{\text{classcial}}(T,\nu)$$

Planck's equation applies the "correction factor"  $\frac{hv}{kT} \frac{1}{e^{hv/kT} - 1}$ 

to the classical expression for the thermal radiation spectrum. *Why*? At temperature *T*, using the Boltzmann distribution, the probability of oscillator energy *nhv* is proportional to  $e^{-nhv/kT}$  and the average energy of the **quantized** thermal oscillators is

$$= \sum_{n=0}^{\infty} P_n nhv = \frac{0 + hv e^{-hv/kT} + 2hv e^{-2hv/kT} + 3hv e^{-3hv/kT} + \cdots}{1 + e^{-hv/kT} + e^{-2hv/kT} + e^{-3hv/kT} + \cdots}$$
$$= \frac{hv}{e^{hv/kT} - 1}$$

# **Applications of Thermal Radiation ?**

Important applications covered next week.

Interesting application: **Vantablack<sup>™</sup>** pigment

The darkest black substance ever synthesized.

Absorbs > 99.9 % of incident light.

Constructed from <u>vertically</u> <u>a</u>ligned <u>nanotubes</u>.

# <u>Vantablack</u>™





a second "violation" of classical physics:

# 2. <u>Heat Capacity of Monatomic Crystals</u>

In the early 1800s, French scientists Dulong and Petit measured heat capacities at constant volume (the change in energy per degree) for different materials.

 $C_{Vm} = (\partial U_m / \partial T)_V$ 

They found  $C_{Vm} \approx 25 \text{ J K}^{-1} \text{ mol}^{-1}$  for monatomic crystals, such as lead and other metals. This is easily understood using classical physics: each atom oscillates in three dimensions (x, y, z) and each of these "degrees of freedom" according to the **equipartition theorem** has on average energy equivalent to kT. T is the temperature and k is Boltzmann's constant (1.38066 × 10<sup>-23</sup> J K<sup>-1</sup>).

Then for one mole  $U_{Vm} \approx N_{Avogadro} \times 3kT = 3RT$  for the molar energy and  $C_{Vm} = [\partial(3RT)/\partial T]_V = 3R = 24.9 \text{ J K}^{-1} \text{ mol}^{-1}$  for the molar heat capacity, in good agreement with Dulong and Petit's experiments.

## But ...

As technology improved, it became possible to measure heat capacities at very low temperatures (approaching absolute zero, T = 0 K). For some reason, unknown at the time, the heat capacities of cold monatomic crystals dropped well below the predicted 3R value. Something was literally:

#### freezing out atomic vibrations at low temperatures Why ? ...

# Einstein's Solution: *Quantized Atomic Vibrations*

Following Planck's success in accounting for blackbody radiation by allowing for only discrete "chunks" of radiation energy, Einstein provided a solution to the heat capacity problem by assuming that the allowed energies of atomic vibrations are not continuous, but can only have discrete *quantized* values:

#### 0, hv, 2hv, 3hv, 4hv, ...

h is Planck's constant, v is the fundamental vibrational frequency.





#### Einstein's calculations:

According to the recently discovered **Boltzmann distribution law**, Einstein reasoned the probability of finding an atom oscillating with energy nhv is proportional to  $e^{-nhv/kT}$  with n = 0, 1, 2, 3, ...

So the average vibrational energy per atom (in three dimensions) is

$$U_{\rm m}(T) = 3 \frac{0 + hv e^{-hv/kT} + 2hv e^{-2hv/kT} + 3hv e^{-3hv/kT} + \cdots}{1 + e^{-hv/kT} + e^{-2hv/kT} + e^{-3hv/kT} + \cdots}$$

After some algebra, the average vibrational energy per mole is

$$U_{\rm m}(T) = 3RT \frac{h\nu}{kT} \frac{1}{{\rm e}^{h\nu/kT} - 1}$$

**Limits?** As an informative exercise, use Einstein's equation to show:

**1.** as  $T \to 0$  (*hv* >> *kT*),  $U_m \to 0$  ("new" quantum behavior)

**2.** as  $T \to \infty$  (*hv* << *kT*),  $U_m \to 3RT$  (classical behavior)

# Einstein's Expression for the Molar Heat Capacity of a Monatomic Crystal

Taking the temperature derivative of  $U_m$  gives

$$C_{Vm} = \frac{\mathrm{d}U_m(T)}{\mathrm{d}T} = 3R \left(\frac{h\nu}{kT}\right)^2 \left(\frac{\mathrm{e}^{h\nu/2kT}}{\mathrm{e}^{h\nu/kT}-1}\right)^2$$

with:  $C_{Vm} \rightarrow 0$  as  $T \rightarrow 0$  (*hv* >> *kT*, all vibrations *frozen* out)

 $C_{Vm} \rightarrow 3R$  as  $T \rightarrow \infty$  ( $hv \ll kT$ , the classical limit)

#### Good agreement with experiment.

And the value of h from the heat capacities agreed with the value from the analysis of blackbody radiation !

closely related ....

# 3. <u>Heat Capacity of Ideal Diatomic Gases</u>

For the heat capacities of diatomic gases, such as  $O_2$  and  $N_2$  in air, **classical mechanics predicts**:

$$C_{Vm} = (\partial U_m / \partial T)_V$$
  
= 7R/2  
= 29.1 J K<sup>-1</sup> mol<sup>-1</sup>

But experimentalists measure:

 $C_{Vm} = 20.8 \text{ J K}^{-1} \text{ mol}^{-1} \approx 5R/2$ 

<u>Why</u>? Just as quantum vibrations of atoms are "frozen out" for cold crystals, diatomic molecules are not actively vibrating, except at very high temperatures.

Equipartition Theory: each energy mode contributes

(classical) RT/2 to the molar energy for a diatomic ideal gas (from Chem 231/232):



	= d(7RT/2)/dT
heat capacity prediction	$C_{Vm} = (\partial U_m / \partial T)_V$
Total energy per mole:	$U_{\rm m} = 7RT/2$
<i>z</i> -axis vibration kinetic ene	rgy RT/2
<i>z</i> -axis vibration potential e	nergy RT/2
<i>x</i> -axis rotation	RT /2
<i>y</i> -axis rotation	RT /2
<i>x</i> -velocity kinetic energy	RT /2
<i>y</i> -velocity kinetic energy	RT /2
<i>z</i> -velocity kinetic energy	RT /2

= 7R/2

7*R*/2 is the classical heat capacity prediction. But experiments give  $C_{Vm} = 5R/2$ . An *R* is missing! Why?

**Equipartition Theory:** each energy mode contributes

(quantum) RT/2 to the molar energy for a diatomic ideal gas (from Chem 231/232):



RT /2 RT /2 RT /2	
RT /2 RT /2	
<del>rgy RT/2</del> hergy RT/2	frozen out
$U_{\rm m} = 5RT/2$	
$C_{\rm Vm} = (\partial U_{\rm m} / \partial T)$	`) <sub>V</sub>
on) = d(5R)	T/2)/d <i>T</i>
	RT/2 $RT/2$ $RT/2$ $RT/2$ $RT/2$ $RT/2$ $RT/2$ $T/2$

= 5R/2

Quantum mechanical vibrational energy levels are too high in energy to be significantly occupied, except at very high temperatures.

# 4. <u>The Photoelectric Effect</u>

Another phenomena of historic interest that "deviated" from classical physics was the photoelectric effect.

**Photoelectric effect**: The ejection of electrons from the surface of a metal by radiation.



kinetic energy of emitted electrons can be measured

The classical picture of light is an oscillating electromagnetic wave. Electrons at the surface oscillate with the changing electric field so violently that they get knocked out or emitted.

This classical picture predicts the kinetic energy of the electrons should increase as the **amplitude** of the radiation (intensity) increases.



## **Curious Experimental Observations:**

- The kinetic energy of the ejected electrons increases with the **frequency** of the light, *not its intensity*.
- No matter how intense the light of a given frequency, the energy of the ejected electrons remains the same.
- Increasing the light intensity increases the number of electrons ejected, but not their kinetic energy.
- For each metal, no electrons are ejected below a threshold frequency of the light specific to that metal.

# **Einstein's Interpretation of the Photoelectric Effect**

Einstein made a major conceptual extension to Planck's concept of quantization, that *radiation exists as small packets of energy,* now called *photons*.

 $E = hv = hc / \lambda$ 

The kinetic energy of the ejected electrons is the energy of the incident photon minus the potential energy (the work function  $\phi$ ) holding the electron in the metal.

 $\frac{1}{2mv^2} = hv - \phi$ 

The slope of the electron kinetic energy plotted against frequency was in good agreement with Planck's h value from studies of blackbody radiation and heat capacities!



The same constant h determined from completely different experiments!



Nobel Prize for the explanation of the photoelectric effect.

### **Measurement of the Photovoltaic Effect**

To measure the maximum kinetic energy of emitted photoelectrons, apply a negative potential to stop all photoelectrons reaching the collector plate.



Apply a negative "stopping" potential to the collector plate to Measure the maximum kinetic energy of the emitted electrons.

## **Photoelectron Data for Sodium Metal**

*Exercise* : show that the slope of the plot is *h* (Planck constant)



Important application of the photoelectric effect (one of many):

#### **Photomultiplier Tubes**

Ultrasensitive radiation detectors. Can detect single photons.



## Some energy magnitudes and units

Using Einstein's relation E = hv between the energy of a photon and its frequency and the relation  $c = \lambda v$  between the frequency, wavelength and speed of light ( $c = 2.99792458 \times 10^8 \text{ m s}^{-1}$ ), the energy of one photon of yellow light (600 nm) is:

 $E = hv = hc/\lambda = 3 \times 10^{-19} \text{ J}$ 

**<u>Electron Volt</u>** When dealing with energies of individual photons, units of electron volts (eV) are commonly used (the energy required to drop one electron through one volt or to raise one proton through one volt):

#### $1 \text{ eV} = 1.602 \text{ x} 10^{-19} \text{ J}$

A photon of yellow light has an energy of about 2 eV.

An  $O_2$  bond energy is about 498 kJ/mol, which is about 5.2 eV.

# 5. <u>Discrete Emission Spectra of the Hydrogen Atom</u> (and other atoms too)

Discrete line spectra of atoms also suggests energy quantization.



Balmer, Rydberg and others *empirically* derived a formula for the hydrogen line spectra:

$$v = \text{constant} \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

 $n_1$ ,  $n_2$  are integers with  $n_2 > n_1$  and the **Rydberg constant** is  $3.28803 \times 10^{15} \text{ s}^{-1}$ 

Hydrogen spectra controlled by two integers - further evidence for some kind of *quantization*.

#### **Bohr Model of the Hydrogen Atom**

The classical picture of an electron and proton cannot explain the discrete ("line") emission spectra.

Bohr came *close*. He *assumed* that the angular momentum of the electrons is quantized with *allowed* values  $\mu Vr = nh/2\pi$ 

 $h/2\pi$ ,  $2(h/2\pi)$ ,  $3(h/2\pi)$ ,  $4(h/2\pi)$ , ...

The centrifugal force on the rotating electron is balanced by the attractive Coulombic force

 $\mu v^2 / r = e^2 / (4\pi \varepsilon_0) r^2$  or  $\mu^2 v^2 r^2 = (nh/2\pi)^2 = [\mu e^2 / (4\pi \varepsilon_0)] r$ 

#### then r is quantized too!

 $\mu = m_e m_p / (m_e + m_p) \approx m_e$  is the reduced mass of the proton and electron, **v** is the electron velocity, **r** is the distance between the proton and electron, and  $\varepsilon_0$  is the vacuum permittivity.

## Angular momentum. What is it?

**Linear momentum** for the motion of a particle in one direction is the product mv of the particle mass m and the particle velocity v.

A particle rotating at a distance r from a fixed center moves through the distance  $2\pi r$  (one circumference) per revolution at the speed  $r\omega$ where the  $\omega$  is the angular velocity defined as the number of radians per second ( $2\pi$  radians per revolution). So  $v = \omega r$ .

The kinetic energy of a particle is  $\frac{1}{2}mv^2 = \frac{1}{2}mr^2\omega^2$  suggesting the analogies:

- $m \leftrightarrow I$  between the mass *m* and moment of inertia  $I = mr^2$
- $\mathbf{v} \leftrightarrow \boldsymbol{\omega}$  between linear and angular velocities
- $mv \leftrightarrow I\omega$  between linear and angular momentum.

Angular momentum =  $I\omega = mr^2 v/r = mvr$ 

Assuming quantized angular momentum, the radii are  $r = n^2 a_0$ where  $a_0$  is the radius of a ground-state hydrogen atom (n = 1):

**Bohr** radius: 
$$a_0 = \varepsilon_0 h^2 / \pi \mu e^2 = 52.9177 \times 10^{-12} \text{ m}$$

Adding the kinetic energy  $\mu v^2/2$  and the potential energy  $-e^2/(4\pi\epsilon_0)r$  from Coulomb's law, Bohr derived the expression  $E_n = -\mu e^4/(8n^2h^2\epsilon_0^2)$  for the quantized energies of a hydrogen atom.

Using the energy change  $\Delta E = E_{n2} - E_{n1} = hv$  for a hydrogen emission line for a transition from  $n_2$  to  $n_1$  gave the frequencies:

$$v = \frac{\mu e^4}{8\varepsilon_o^2 h^3} \left(\frac{1}{n_1^2} - \frac{1}{n_2^2}\right)$$

Except for the assumption of angular momentum quantization, the Rydberg equation was derived with classical mechanics.

$\mu$ = reduced mass of electron a	and proton = $m_e m_p / (m_e + m_p) \approx m_e$
$m_{\rm e}$ = electron mass	(9.10939 x 10 <sup>-31</sup> kg)
$m_{\rm p}$ = proton mass	(1.67262 x 10 <sup>-27</sup> kg)
e = elementary charge	(1.602177 x 10 <sup>-19</sup> C)
$\varepsilon_0 = vacuum permittivity$	(8.85419 x 10 <sup>-12</sup> J C <sup>2</sup> m <sup>-1</sup> )
<i>h</i> = Planck's constant	(6.62608 x 10 <sup>-34</sup> J s)

Using  $c = \lambda v$  and expressing frequencies in wavenumbers  $(\lambda^{-1} = v/c)$ : Bohr predicted (in amazing agreement with experiment):  $v/c = (109,677 \text{ cm}^{-1})(1/n_1^2 - 1/n_2^2)$  Hydrogen Emission Spectra Experimentalists measured:  $v/c = (109,677.576 + \text{ or} - 0.012 \text{ cm}^{-1})(1/n_1^2 - 1/n_2^2)$ 

Lyman series:

 $n_1 = 1, n_2 = 2, 3, 4, ...$  (ultraviolet)  $n_1 = 2, n_2 = 3, 4, 5, ...$  (visible)  $n_1 = 3, n_2 = 4, 5, 6, ...$  (infrared)  $n_1 = 4, n_2 = 5, 6, 7, ...$  (infrared)

#### de Broglie's Wave Nature of Matter

#### A <u>conceptual break through</u> came from de Broglie in 1923.

By then the dual wave-particle nature of light was accepted by most leading physicists.

de Broglie extended and *quantified* the wave-particle duality of light to all particles (electrons, protons, baseballs).

Using special relativity, Einstein had derived an expression for the momentum of a photon (even though it has no mass). Using a similar line of reasoning, de Broglie argued that the *wavelength*  $\lambda$  of any *particle is related to its momentum* p = mv as follows:

$$\lambda = \frac{h}{mv} = \frac{h}{p}$$

*h* is Planck's constant, *again!* 

For photons, which travel at the speed of light (c), combining  $E = hv = hc/\lambda$ 

and

 $E = mc^2$  (from the theory of relativity)

gives

 $\lambda = h/mc$ 

de Broglie extension to particles of mass *m* moving at velocity v:

 $\lambda = h/mv = h/p$ 

And de Broglie's concept of the wave nature of particles explained Bohr's assumption that the angular momentum of electrons in H atoms are quantized! For an electron in a stable orbit around a proton, the orbit circumference is an integer multiple of the electron wavelength  $(2\pi r = n\lambda)$  which gives quantized angular momentum:

 $mvr = nh/2\pi$  (Bohr's assumption)

In 1925, experimental verification of the particle wave hypothesis came with electron diffraction experiments.



X-ray and electron diffraction patterns through aluminum foil



Louis de Broglie, Nobel Prize for the discovery of the wave nature of electrons.

Today, the wave property of matter is used routinely:

- electron microscopes (electron microscopes have shorter wavelengths and higher resolution than optical microscopes)
- neutron diffraction

Wave-particle duality of matter is significant only for very small particle masses.

h 2 p mV

Particle	Mass (kg)	Speed (m/s)	Wavelength (1 pm = 0.001 nm)
Electron accelerated through 100 V	9.11x10 <sup>-31</sup>	5.9x10 <sup>6</sup>	<b>120 pm</b> (atomic and molecular distances)
Alpha particle ejected from radium	6.68x10 <sup>-27</sup>	1.5x10 <sup>7</sup>	6.6x10 <sup>-3</sup> pm (smaller than an atom)
Bullet	1.9x10 <sup>-3</sup>	3.2x10 <sup>2</sup>	1.1x10 <sup>-21</sup> pm (much smaller than a nucleus)

For macroscopic bodies, de Broglie wavelengths are completely undetectable and of no practical consequence.

#### Recall the important question about the existence of atoms:

Why doesn't the electron spiral inwards to merge with the proton, forming a **subatomic particle** (diameter << 0.10 nm for H atoms)?

#### Hydrogen Atom

electron and a proton

positive and negative charges attract:

electron proton 0.0002 nm? ( subatomic particle)

#### Answer: electrons are waves !



#### Huge Significance:

- formation of atomic (1s, 2s, 2p, ...) and molecular orbitals as "standing" waves
- standing (resonance) waves have fixed energy  $\rightarrow$  quantization of energy
- existence of atoms and molecules (and all of chemistry !)

\*At <u>very high</u> energies (millions of eV), electron wavelengths are much smaller, allowing electrons and protons to merge, as in the formation of neutron stars.

#### **Wave-Particle Duality**

**Wave-particle duality** refers to the fact that both light and matter can exhibit particle-like behavior or wave-like behavior depending on how we observe them.

In other words, their behavior depends on the nature of the experiment.

- photons can behave like particles in a photo-electric experiment
- electrons and other particles can exhibit a wave-like diffraction patterns



The density of an atomic cloud as the temperature is decreased (left to right): wave-like behavior

#### Cooling Technology

Atoms are first cooled by lasers, then confined by a magnetic trap.

Evaporative cooling is the last step.

Atoms can all be described by a single wave function!

# **2001 Nobel Prize in Physics**

#### Eric A. Cornell

JILA and National Institute of Standards and Technology (NIST), Boulder, Colorado, USA,

#### **Wolfgang Ketterle**

Massachusetts Institute of Technology (MIT), Cambridge, Massachusetts, USA, and

#### Carl E. Wieman

JILA and University of Colorado, Boulder, Colorado, USA,

"for the achievement of **Bose-Einstein condensation** in dilute gases of alkali atoms, and for early fundamental studies of the properties of the condensates".





Bose

Einstein

Cooling rubidium atoms to less than 170 billionths of a degree above absolute zero caused the individual atoms to condense into a "superatom" behaving as a single entity







a few thousand atoms

Some atoms in a BEC condensate

Some very cold atoms

mid 1920s

The wave-particle duality of both light and matter leads to some very awkward results.

Consider the measurement of the position of an electron.

If we want to locate the electron within a distance  $\Delta x$ , we must use spatial **resolution** less than  $\Delta x$ .

One way to achieve this is to use light of wavelength  $\lambda \approx \Delta x$ .

For us to 'see' the electron, the photon must interact with the electron.

But the photon has a momentum associated with it.  $p = \frac{n}{\lambda}$ 

# Thus, the very act of observing the electron leads to a change in its momentum.

Classical mechanics: no limitations on the accuracy of measurements of position, velocity, momentum, ... of particles

But wait ! From quantum mechanics, particles are actually waves !

 $\Delta x \Delta \frac{1}{\lambda} \geq \frac{1}{4\pi}$ 

For wave motion (from Fourier integral methods):

#### **Example:**

<u>A Monochromatic Wave ( one fixed wavelength  $\lambda$  ) Moving at Velocity v</u>

 $\Delta \frac{1}{\lambda} = 0$  so  $\Delta x = \infty$  It's everywhere !

$$\Delta x \, \Delta \frac{1}{\lambda} \geq \frac{1}{4\pi}$$



As the number N of different wavelengths increases,  $\Delta(1/\lambda)$  increases, so  $\Delta x$  decreases, causing the wave packet to be more localized in space, behaving more like a particle.

$$\Delta x \, \Delta \frac{1}{\lambda} \geq \frac{1}{4\pi}$$

**Delta Function (superposition of**  $\infty$  **different waves)** 

$$\psi(x,t) = \sum_{i=1}^{\infty} A_i \sin\left[\frac{2\pi}{\lambda_i}(x - vt)\right]$$

In the limit of an infinite number of superimposed waves,  $\Delta(1/\lambda) = \infty$  and  $\Delta x = 0$ , so the "particle" is exactly located.

$$\Delta x \, \Delta \frac{1}{\lambda} \geq \frac{1}{4\pi}$$

wave velocity = frequency  $\times$  wavelength

$$v = v \lambda$$

divide and multiply by V:

$$\Delta \frac{x}{v} \Delta \frac{v}{\lambda} \geq \frac{1}{4\pi} \longrightarrow \qquad \Delta t \Delta v \geq \frac{1}{4\pi}$$

#### **Application to spectroscopy:**

Transitions involving long-lived states (large  $\Delta t$ ) absorb or emit nearly monochromatic radiation.

Short-lived states produce broader ranges of frequencies.

$$\Delta x \, \Delta \frac{1}{\lambda} \geq \frac{1}{4\pi}$$

Multiply by Planck constant h:

$$\Delta x \, \Delta \frac{h}{\lambda} \geq \frac{h}{4\pi}$$

Recall that  $h/\lambda$  is the momentum p (the de Broglie relation):

$$\Delta x \, \Delta p \geq \frac{h}{4\pi}$$
  
In written as: 
$$\Delta x \, \Delta p \geq \frac{\hbar}{2}$$

Usually

"h-bar" is the Planck constant divided by  $2\pi$ :

$$\hbar = \frac{h}{2\pi}$$

Developing this idea more fully, Werner Heisenberg showed that it is not possible to simultaneously determine the **exact** position x and the **exact** velocity v of a particle at the same time.

The smaller the uncertainty  $\Delta x$  in the position of a particle, the larger the uncertainty in the particle's momentum  $\Delta p$ , and *vice versa*.

$$\Delta x \Delta (mv) \geq \frac{\hbar}{2}$$
  
Heisenberg Uncertainty Principle

$$\hbar = \frac{h}{2\pi}$$

The uncertainty principle is incompatible with the deterministic classical picture because we cannot specify exactly a particle's position and momentum simultaneously. We can only talk about *probabilities*.

But the uncertainty principle is important only on *microscopic scales*.

 $h = 6.626 \times 10^{-34} \text{ J s}$  (small!)

*e.g.* The uncertainty in the position of a baseball (145 g) thrown at 90 mph (40 m/s) if we measure the momentum (p = mv) to a millionth of 1.0% (9 x 10<sup>-8</sup> mph):

$$\Delta x \Delta p \ge \frac{\hbar}{2} = 5x10^{-35}$$

 $\Delta p = 5.6 \times 10^{-8} \text{ kg m/s}$ 

 $\Delta x = 9.4 \text{ x10}^{-28} \text{ m}$ 

(negligible, less than the radius of atomic nuclei )

*e.g.* The uncertainty in the momentum if we locate an electron within an atom so that the uncertainty in its position is 50 pm. (Bohr radius)

If 
$$\Delta x = 50 \times 10^{-12} \text{ m}$$
  
Then  $\Delta p = m \Delta v = 1.3 \times 10^{-23}$   $\longrightarrow \Delta v = 1 \times 10^7 \text{ m/s}$   
kg m/s  
Relatively large!

## Historical Development of Quantum Mechanics

The stage was set for the development of a new theory to describe the microscopic world of electrons and nuclei.

- classical mechanics combined with ideas of quantization can reproduce experimental observations.
- quantization of energy states with Planck's constant
- wave-particle duality of light and matter
- uncertainty principle starts to hint at probabilities (not classical determinism)
- Boltzmann's law is used to calculate the probabilities that atoms and molecules are in allowed energy states

# **Modern Physics:**

- theory of relativity
- quantum mechanics

**Theory of Relativity** : Developed by Einstein in 1905, extended classical mechanics to high velocities and astronomical distances.

**Quantum Mechanics**: developed over decades by many scientists. Deals with the microscopic at the level of atoms, electrons and smaller.

Quantum Mechanics has had a profound effect on our understanding of chemistry. For example, covalent bonds cannot be explained by classical physics. There is a subdiscipline of chemistry and quantum mechanics called 'quantum chemistry'.

## **Quantum Mechanics**

Modern quantum mechanics was "discovered" in 1925.

In 1925, during Christmas holidays, Schrödinger developed the Wave Equation of quantum mechanics, starting from de Broglie's idea that particles are waves

Schrödinger's formulation of quantum mechanics, known as *Wave Mechanics,* replaced the "old" quantum mechanics developed by Bohr.



In quantum chemistry, Schrödinger's formulation of quantum mechanics is most frequently used. In other applications, especially nuclear physics, Heisenberg's and other formulations are used.

There are other notable developments of quantum mechanics:

- 1925 Heisenberg's Matrix Quantum Mechanics (1932 Nobel Prize)
- 1926 Schrödinger showed that Matrix and Wave Mechanics are equivalent

Schrödinger's and Heisenberg's treatments of quantum mechanics do not include Einstein's theory of relativity.

1929 – Dirac – *Relativistic Quantum Mechanics* (1933 Nobel Prize)

Complicated, and not widely used in chemistry. Includes antimatter. **Spin** is introduced as a natural variable, just like time or position, but nothing is spinning!

Spin is included empirically in quantum chemistry.

1941 – Feynman (1965 Nobel Prize) – Path Integral Formulation of Quantum Mechanics *"I think it's safe to say that nobody really understands quantum mechanics."* Richard Feynman, 1965 Nobel Physics Prize for research on quantum electrodynamics.



1. Introduction – Why is Quantum Mechanics Important for Understanding Chemistry?

Important topics and terminology:

thermal radiation (*aka* blackbody radiation, *aka* cavity radiation)

Planck's radiation law Planck constant h quantization **Boltzmann distribution law** h bar  $(h/2\pi)$  $P_i \propto \exp(-E_i/kT)$ crystal vibrations low-temperature heat capacities photoelectric effect Bohr model of the H atom linear momentum angular momentum Bohr radius  $a_0$ continuous spectra line spectra Rydberg constant de Broglie wavelength *h/p* de Broglie relation electron diffraction wave/particle duality Heisenberg uncertainty principle