# Temperature-dependent absorption cross-sections of HCFC-142b

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# 8 Abstract

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Following the recent detection of HCFC-142b (1-chloro-1,1-diffuoroethane) from space, laboratory infrared absorption cross-section spectra of this molecule 10 in a pure vapour phase have been recorded in the  $650-3500 \text{ cm}^{-1}$  spectral re-11 gion using Fourier transform spectroscopy. The spectra have been recorded 12 at a resolution of  $0.02 \text{ cm}^{-1}$  and a range of temperatures from 223 to 283 13 K. The resulting data show good agreement with the harmonic frequencies 14 and intensities calculated using density functional theory as well as with the 15 integrated absorption intensities of the spectral bands available in the litera-16 ture. The new cross-sections will allow more accurate retrieval of atmospheric 17 HCFC-142b concentrations using infrared spectroscopic techniques. 18 Key words: 19 Hydrochlorofluorocarbons, R-142b, cross-sections, mid-infrared, FTIR, 20 gas phase, density functional theory, vibrational wavenumber, band 21

22 strength

Preprint submitted to Journal of Quantitative Spectroscopy and Radiative TransferOctober 15, 2009

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#### 23 1. Introduction

Hydrochlorofluorocarbons (HCFCs) are temporary substitutes for chlo-24 rofluorocarbons (CFCs). They were introduced after the phase-out of the 25 latter by the Montreal Protocol and its subsequent amendments. Unlike 26 CFCs, which are mainly destroyed by solar ultraviolet radiation in the strato-27 sphere, the HCFCs, which contain one or more hydrogen atoms, can react 28 with OH radicals in the troposphere to create HF and  $CO_2$ . Therefore, as 29 less chlorine and fluorine are transported to the stratosphere, the ozone de-30 pletion potentials of HCFCs are substantially weaker than those of CFCs. 31 However, because of their C-Cl and C-F bonds, these molecules still have 32 large absorption cross-sections in the atmospheric window region (8-12  $\mu$ m), 33 which give them strong global warming potentials. 34

HCFC-142b (1-chloro-1,1-diffuoroethane) is a colorless gas at ambient 35 pressure. It is mainly used as a chemical intermediate to produce fluoropoly-36 mers, as a blowing agent for expanded polystyrene and as a component of 37 refrigerant fluids. Because of its high vapour pressure and low vapour solu-38 bility, HCFC-142b partitions mostly in the atmosphere. Today, this is the 39 third most abundant hydrochlorofluorocarbon after HCFC-22 and HCFC141-40 b. The atmospheric lifetime of HCFC-142b is 17.9 years  $\pm 24\%$  [1]. Its ozone 41 depletion potential has been estimated at 0.07 while its global warming po-42 tential is evaluated to be  $2310 \pm 810$  for a horizon of 100 years [1]. 43

The volume mixing ratio of HCFC-142b has been rising steadily in the atmosphere since the beginning of the 1990's, exceeding 20 ppt in 2008 at 13-16 km with an average positive trend of more than 5% per year [2]. The atmospheric concentration of HCFC-142b is now high enough to allow its

detection from space [2, 3]. However, the accuracy of the retrieval is limited 48 by the lack of spectroscopic knowledge of this molecule. Only three cross-49 section spectra corresponding to the temperatures 253 K, 270 K and 287 50 K are currently available in the HITRAN 2008 database [4]. Measurements 51 of temperature-dependent infrared cross-sections of HCFC-142b vapour at 52 low temperature were reported in the 1990's [5, 6, 7] but discrepancies exist 53 between these studies. No new laboratory data at low temperature have been 54 published on this molecule since 1995. 55

As a result, uncertainties on the spectroscopic parameters of HCFC-142b are seen as one of the main sources of error for atmospheric retrievals [3]. New laboratory measurements are therefore crucial to enable accurate observations of the spatial and temporal variation of this hydrochlorofluorocarbon in the atmosphere.

The purpose of this study is to provide new infrared (IR) cross-sections of HCFC-142b at relatively high resolution (0.02 cm<sup>-1</sup>) and at a range of relevant atmospheric temperatures (from 223 to 283 K). The resulting experimental data are compared to theoretical calculations performed using density functional theory and to previously published values.

# 66 2. Experimental setup

Experimental data are obtained using Fourier transform infrared (FTIR) absorption spectroscopy. The Fourier transform spectrometer (FTS) is a Bomem DA8.002 equipped with a KBr beamsplitter and operating with a Globar source.

The gas sample (SynQuest Labs, purity > 98%) is contained in a stainless

steel cell positioned between the FTS and a liquid nitrogen-cooled mercury
cadmium telluride (MCT) detector. A vacuum system equipped with a Varian Turbo-V V250 turbo molecular pump allows the cell to be evacuated to
about 1.10<sup>-6</sup> torr. The cell pressure is measured using 10 Torr and 1000 Torr
MKS baratron pressure gauges simultaneously.

HCFC-142b presents very intense lines in the atmospheric window region.
In order to avoid saturation effects while working at relevant pressures, a
stainless steel cell with an optical path of 3.17 cm was constructed. ZnSe
windows were sealed to the cell with indium O-rings to prevent leakage at
low temperature. The windows are maintained in place by stainless steel
flanges supporting teflon rings.

The cooling is achieved by a Neslab ULT-80 chiller sending the coolant (Syltherm XLT) to a copper tube surrounding the cell. The copper tube is soldered to the cell and covered by thermally conductive epoxy to provide temperature homogeneity. The cell temperature is measured by a thermocouple directly inserted inside the cell. The temperature readout accuracy during experiments is typically  $\pm 0.1^{\circ}$ C.

Many sources of errors and artefacts can affect an FTIR spectrum, such as spectral aliasing, dynamic alignment error, blackbody emission from the source aperture, non-linearity of the MCT detector in the mid-infrared, etc. The spectral aliasing is reduced by the acquisition of a recursive interferogram signal produced by a monochromatic reference source (He-Ne laser) and a digitization resolution of 16 binary bits. The dynamic alignment is corrected in the Bomem DA8 spectrometer by a patented electronic tilt correction. The r.m.s. angular deviation from optimum alignment is less than 10<sup>-6</sup> radians <sup>97</sup> in normal laboratory conditions (data provided by the manufacturer).

The blackbody emission from the source aperture comes from the warm-98 ing of the iris by the light source. The warm annulus around the aperture 99 acts as a second infrared source emitting off-axis thermal radiation. This 100 leads to distortion of the signal shape and intensities as the aperture size 101 decreases. To reduce this warm aperture artefact, a second iris is inserted 102 between the DA8 spectrometer and the MCT detector at the focal point of 103 the beam [8]. After this installation, no variation of the cross-section signal 104 with the aperture size has been observed. 105

The non-linearity of the MCT detector, corresponding to the non-linearity 106 between the measured signal and the photon flux, essentially affects the cen-107 tral fringe of the interferogram, which leads to a zero-level offset in the ab-108 sorption spectrum. Correction of non-linearity is applied on the raw inter-109 ferograms before the phase correction and the Fourier transformation. We 110 adjusted the response of the detector to the light intensity by a curve whose 111 only unknown parameter is an empirically chosen DC output value [9]. This 112 parameter is adjusted by a minimization of the transmission in the spectral 113 region below the cut-off wavenumber. The residual baseline offset is sub-114 tracted from the spectra after phase correction and Fourier transformation. 115

# 116 **3. Data analysis**

For each temperature, a series of spectra were recorded at a minimum of six different pressures between 2 and 12 torr. Each Pressure-Temperature (P-T) spectrum is composed of at least 200 unapodised scans. The resulting spectrum, for each temperature, is a composite spectrum from the P-T spec-

trum sets extrapolated to the zero-torr limit. A 200-scan primary baseline 121 spectrum with an empty cell was recorded at each temperature. Secondary 122 control baselines of at least 20 scans were also taken before and after each 123 sample measurement to account for the small intensity variations that can 124 occur during the long periods (typically several hours) of acquisitions. If 125 necessary, the primary baseline is adjusted to the secondary baselines by a 126 polynomial regression prior to the division of each sample spectrum by the 127 adjusted baseline spectrum. 128

The cross-section,  $\sigma(\nu)$ , in cm<sup>2</sup>/molecule is calculated for each P-T set using the well-known Beer-Lambert law:

$$I(\nu) = I_0(\nu) e^{-\chi(\nu)}$$
(1)

<sup>131</sup> with optical depth:

$$\chi(\nu) = \sigma(\nu) \frac{PT_0}{TP_0} N_L L \tag{2}$$

where  $\nu$  is the wavenumber (cm<sup>-1</sup>);  $I_0(\nu)$ , the light intensity passing through the empty cell (baseline);  $I(\nu)$ , the light intensity passing through the sample gas cell;  $N_L$  is the Loschmidt's constant (2.6868x10<sup>19</sup> molecules/cm<sup>3</sup>); L, the length of the cell (cm); and  $P_0$  and  $T_0$ , the standard conditions for pressure and temperature.

To prevent saturation effects in an optically thick medium while keeping a good signal-to-noise ratio at every wavenumber, the points corresponding to optically thick ( $\chi(\nu) > 1.1$ ) or optically thin ( $\chi(\nu) < 0.1$ ) conditions are eliminated. This way, a linear behaviour is obtained for strong absorption bands from the low-pressure measurements while the weak absorption features are represented by the high-pressure measurements. The shapes of the ro-vibrational transition lines, as well as the peaks of the Q-branches, are pressure-dependent due to collisional broadening. The crosssection for each wavenumber is obtained at the zero-torr limit by the linear least-square fit of the experimental apparent cross-section of the remaining values  $(0.1 < \chi(\nu) < 1.1)$  versus the pressure.

The errors on optical path length, temperature readout, pressure, baseline 148 drift, and sample purity have been evaluated to account for less than  $\pm 3\%$ . 149 The other sources of error in spectral measurements come from the residual 150 MCT non-linearity, a possible residual baseline drift and the errors induced 151 by the data reduction. For each wavenumber, we calculated the standard 152 deviation  $\Delta$  between the linear fit and the apparent cross-section values as 153 a function of pressure. The error for each wavenumber is chosen at the 95%154 confidence limit  $(2\Delta)$ . 155

#### 156 4. Results

<sup>157</sup> HCFC-142b is a near-prolate asymmetric top molecule belonging to the <sup>158</sup> Cs symmetry group (Figure 1). As a consequence, it has 18 fundamental <sup>159</sup> vibration modes (11 with A' symmetry and 7 with A" symmetry). Twelve <sup>160</sup> of those fundamental modes have active absorption bands in the spectral <sup>161</sup> window region between 650 and 3500 cm<sup>-1</sup>.

All the bands present a sharp Q-branch surrounded by less intense Pand R-branch structures (Figure 2). No significant variation of the integrated band strength with temperature has been observed from 223 to 283 K for any of the bands (Figure 3). Therefore, we can confirm that the superimposed hot bands and combination bands of HCFC-142b do not play a significant <sup>167</sup> role in our range of temperatures.

The shape of the bands is strongly dependent on the temperature. The *Q*-branches become sharper as the temperature decreases due to the change of relative populations of the rotational states of the vibrational bands. This leads to a significant increase in the cross-sections at the band centers with decreasing temperature (Figure 4).

The uncertainty on the integrated band strength is stable with temperature for all bands with the exception of the  $\nu_8$  band. The increased uncertainty on the integrated strength of the  $\nu_8$  band as the temperature decreases is likely due to the intense sharpening of the  $\nu_8$  *Q*-branch which worsens the data reduction error on the main peak.

Weak features of magnitude lower than  $10^{-20}$  cm<sup>2</sup>/molecule, which can be attributed either to overtones and combination band or sample impurity, are present at all temperatures in the spectral region between 1500-2800 cm<sup>-1</sup>.

#### <sup>181</sup> 5. Data Validation

# 5.1. Comparison with theoretical calculation of vibrational modes and inten sities

The unconstrained geometric optimizations and harmonic vibrational frequencies of HCFC-142b are calculated using density functional theory (DFT) with Gaussian 03 [10].

Previous ab-initio calculations for HCFC-142b have been carried out at the RHF/6-31\* [11] and MP2/6-31G\*\* [12] levels . The present DFT calculations are performed using the Becke's three-parameter functionals for the exchange and the Lee-Yang-Parr and Perdew-Wang non-local functionals for <sup>191</sup> correlation (respectively B3LYP and B3PW91). For each functional set, two
<sup>192</sup> basis sets are used: the Pople's type valence triple-zeta basis supplemented by
<sup>193</sup> multiple polarization and diffuse functions (6-311++(3df,3pd)) and the aug<sup>194</sup> mented correlation-consistent polarized quadruple-zeta (aug-cc-pVQZ) basis.

We observe a very good agreement between the geometries from the four 195 levels of theory. The choice of the correlation functionals and the size of 196 the basis sets have only a weak influence on the geometry. The harmonic 197 wavenumber positions present little sensitivity to basis sets, with an average 198 difference below 0.2%, and to the correlation functionals, with an average 199 difference around 0.7%. The variations of intensity with basis sets are also 200 weak in both cases, with the exception of the  $\nu_{10}$  band, corresponding to 201 C-Cl stretching. It may be noted that the inclusion of diffuse functions 202 on hydrogen atoms has almost no influence on the geometry and harmonic 203 frequencies of HCFC-142b. However, the choice of correlation functionals 204 does have a strong influence on the intensities, with an average difference 205 between LYP and PW91 above 12% and a maximum difference between 206 25-30% for the  $\nu_4$  and  $\nu_{15}$  bands. Comparison with experimental data is 207 therefore essential to test the validity of the methods. 208

Table 1 presents the assignment, the calculated harmonic wavenumber and the integrated strength of the fundamentals modes. The strongest bands correspond to C-C, C-Cl and C-F stretching vibrations. All of them fall in the 800-1250 cm<sup>-1</sup> atmospheric window region, which explains the high global warming potential of this molecule.

The weak fundamental C-H stretching vibrations fall in the 3000 cm<sup>-1</sup> spectral region. The superimposed CH<sub>3</sub> deformation vibrations lie in the

<sup>216</sup> 1400 cm<sup>-1</sup> region. Only the  $\nu_4$  mode presents a medium intensity in this <sup>217</sup> window.

The theoretical values of the harmonic frequencies have been compared 218 to the experimental data (Figure 5). The experimental line center is chosen 219 at the barycentre of the Q-band. All four levels of theory give excellent re-220 sults after an average linear scaling:  $\nu_{exp} = (1.06 \pm 0.01)\nu_{theo} - (59.3 \pm 7.7),$ 221 with  $\nu_{exp}$  and  $\nu_{theo}$ , the respective experimental and theoretical harmonic 222 wavenumber values in  $\rm cm^{-1}$ . The theoretical values for line intensity in the 223 mid-IR spectral windows also fit relatively well the data after linear scal-224 ing with the experimental integrated absorption strength. We obtained an 225 average  $I_{exp} = (1.01 \pm 0.05)I_{theo} + (0.22 \pm 0.14) \times 10^{-17}$  for B3LYP and 226  $I_{exp} = (1.10 \pm 0.02)I_{theo} + (0.08 \pm 0.05) \times 10^{-17}$  for B3PW91, with  $I_{exp}$  and 227  $I_{theo}$  the respective experimental and theoretical integrated band strengths 228 in cm/molecule. However, B3PW91 gives the best correlation with an  $\mathbb{R}^2$  of 229 0.999 for both basic sets. 230

### 231 5.2. Comparison with published values

The integrated band intensities of the HCFC-142b cross-section spec-232 tra are compared with data available in the literature (Table 2). The inte-233 grated intensities of the four absorption bands  $\nu_8$ ,  $(\nu_7, \nu_{15})$ ,  $(\nu_6, \nu_5, \nu_{14})$ , and 234  $(\nu_4, \nu_3, \nu_{13})$  between 650 and 1500 cm<sup>-1</sup> are reported. No previous values 235 for the  $\nu_1, \nu_2, \nu_{12}$  bands around 3000 cm<sup>-1</sup> have been found in literature. 236 At most temperatures, our integrated absorption band strengths show good 237 agreement with earlier measurements for a pure vapour [5, 6], as well as for 238  $N_2$ -broadened vapour [7, 13]. No constant relative difference is observed with 239 any of the data sets, which indicates that there were no significant systematic 240

<sup>241</sup> errors in the measurements.

The largest discrepancies between authors appears at 233 K. A disparity close to 20% for the  $\nu_6$ ,  $\nu_5$ ,  $\nu_{14}$  absorption bands exists between Newnham and Ballard [5] and Cappellani and Restelli [7]. As previously mentioned, our integrated band strengths do not exhibit trends and our values at 233 K, midway between the values of these two studies, remain consistent with our results at other temperatures.

#### 248 6. Conclusions

Absorption cross-sections of HCFC-142b at a spectral resolution of 0.02 249  $\rm cm^{-1}$  have been recorded in the mid-infrared between 650 and 3500  $\rm cm^{-1}$ 250 at seven different temperatures (223, 233, 243, 253, 263, 273, and 283 K). 251 The main sources of errors in the measurements have been investigated and 252 accounted for in the final results. The integrated intensities of the harmonic 253 wavenumber bands show good agreement with the data available in the liter-254 ature and with theoretical calculations using DTF. The availability of cross-255 section spectra at 10-K steps between 223 and 283 K should, therefore, allow 256 the reduction of uncertainties on HCFC-142b volume mixing ratio retrieval 257 from space missions. The cross-sections are provided online in supplementary 258 data files and are also available from the corresponding author. 259

# <sup>260</sup> 7. Acknowledgments

This work was supported by the Canadian Space Agency (CSA) and the Natural Sciences and Engineering Research Council of Canada (NSERC). We thank Paul Chen for technical support, and Prof. James R. Drummond and the NSERC Industrial Research Chair in Atmospheric Remote Sounding
from Space (sponsored by COMDEV, Bomem, Environment Canada, CSA,
and NSERC) for the use of the Bomem DA8 Fourier transform spectrometer.
We also thank Dr. Stella M.L. Melo from the CSA for helpful discussions.

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Figure 1: Geometrical structure of HCFC-142b.

Figure 2: Survey spectrum of HCFC-142b in the mid-infrared at 263 K.

Figure 3: Variation of the integrated strength of the five main absorption bands of HCFC-142b in the mid-infrared.

Figure 4: Cross-section of the P- and Q-branches of the  $\nu_{15}$  band of HCFC-142b at seven different temperatures.

Figure 5: Comparison between experimental and theoretical harmonic frequencies of HCFC-142b for the four levels of theory.

Figure 6: Comparison between experimental and theoretical band strengths of HCFC-142b for the four levels of theory.

ble 1: Theoretical harmonic wavenumbers and intensities	he fundamental modes of HCFC-142b. Wavenumbers are	${\rm cm}^{-1},$ bands strengths are in $10^{-17}$ cm/molecule. (a):6-	++, (b):Aug-cc-pvqz. st:stretch, def: deformation.
whe 1: Theoretical harmon	the fundamental modes of	$cm^{-1}$ , bands strengths ar	++, (b):Aug-cc-pvqz. st:

				,					
Label	Assignment		Waven	umber		щ	3and st	trength	_
		B3L	AP	B3P	W91	B3I	MD	B3P	W91
		(a)	(q)	(a)	(q)	(a)	(q)	(a)	(q)
$\mathbf{A}^{\prime}$									
$   \nu_1 $	$CH_3 sym st.$	3141.36	3136.41	3157.78	3152.34	0.06	0.06	0.05	0.05
$\nu_2$	$CH_3 sym st.$	3063.13	3059.54	3071.15	3066.82	0.03	0.03	0.03	0.03
$\nu_3$	CH <sub>3</sub> sym def.	1483.52	1483.83	1474.41	1474.57	0.01	0.01	0.01	0.01
$ u_4 $	$CH_3 sym def.$	1421.65	1422.37	1412.26	1412.58	0.46	0.47	0.62	0.63
$\nu_5$	$CF_2$ sym st.,	1217.65	1216.27	1231.51	1229.66	2.00	1.99	1.93	1.94
	C-C st.								
$ u_6 $	C-Cl st.,	1129.57	1129.39	1128.24	1127.67	2.72	2.67	3.01	2.95
	$CF_2$ sym. def.								
V7	$CF_2$ sym. st.	892.16	889.67	903.03	900.49	2.10	2.14	1.83	1.88
$\nu_8$	C-C-Cl def.	669.26	668.25	678.87	677.57	0.98	0.98	0.94	0.94

	Tab	ole 1 – co	ntinued	from pre	vious pag	ge			
Label	Assignment		Waven	umber		Η	3and st	trengtl	I
		B3L	$\rm AP$	B3P	W91	B3I	AP	B3P	W91
		(a)	(q)	(a)	(q)	(a)	(q)	(a)	(q)
$\nu_9$	$CF_2$ scissor	538.74	538.32	543.53	543.01	0.22	0.22	0.21	0.21
$ u_{10} $	C-Cl st.	424.24	421.51	431.71	428.52	0.05	0.05	0.04	0.04
$ u_{11} $	$CCIF_2-CH_3 rock$	301.07	299.86	301.40	299.99	0.02	0.02	0.02	0.02
$\mathbf{A}$ "									
$ u_{12} $	$CH_2$ asym. st.	3157.99	3153.42	3173.96	3169.17	0.03	0.03	0.02	0.02
$ u_{13} $	CH <sub>3</sub> asym. def	1480.50	1480.27	1471.51	1470.95	0.06	0.06	0.06	0.06
$ u_{14} $	$CF_2$ asym. st.	1185.37	1183.27	1194.77	1192.94	2.36	2.33	2.64	2.61
$ u_{15} $	$CF_2$ asym. st.	963.01	961.31	968.74	967.87	1.20	1.2	0.93	0.97
$ u_{16} $	$\operatorname{CF}_2\operatorname{rock}$	430.36	429.96	431.12	430.89	0.00	0.00	0.00	0.00
$\nu_{17}$	CCIF <sub>2</sub> -CH <sub>3</sub> twist	330.49	329.79	332.15	331.29	0.01	0.01	0.01	0.01
$ u_{18} $	torsion	236.96	235.49	239.07	238.32	0.00	0.00	0.00	0.00

Table 2: Comparison of the experimental infrared integrated band strengths ( $x10^{-17}$  cm/molecule) with published values when available. (a) The first two temperature sets for Ref. [13] are at 297 and 277 K. (b) The first temperature set for Ref. [5] is at 296 K. (c) The first two temperature sets for Ref. [6] are at 287 and 270 K.

The state	T 1 1		Integra	ated band str	ength	
Iemperature	Label	This work	Ref. $[13]^{a}$	Ref. [5]	Ref. $[6]^{b}$	Ref. [7]
	$\nu_8$		0.74	$0.71 \pm 0.01$		0.74
909 V	$\nu_7,\nu_{15}$		2.53	$2.49 \pm 0.07$		2.38
293 K	$\nu_6,\nu_5,\nu_{14}$		7.06	$7.34 \pm 0.10$		6.94
	$\nu_4,\nu_3,\nu_{13}$		0.64	$0.58 \pm 0.04$		0.60
	$ u_8$	$0.69 \pm 0.02$			$0.75 \pm 0.05$	
909 IZ	$\nu_7,\nu_{15}$	$2.41 \pm 0.07$			$2.58 \pm 0.14$	
283 K	$\nu_6,\nu_5,\nu_{14}$	$6.99 \pm 0.19$			$7.19 \pm 0.20$	
	$ u_4, \nu_3, \nu_{13}$	$0.57 \pm 0.02$			$0.61 \pm 0.09$	
	$\nu_2,\nu_1,\nu_{12}$	$0.13 \pm 0.01$				
	$\nu_8$	$0.70 \pm 0.03$	0.74	$0.70 \pm 0.04$	$0.69 \pm 0.05$	0.75
273 K	$\nu_7,\nu_{15}$	$2.42 \pm 0.07$	2.50	$2.56 \pm 0.07$	$2.51 \pm 0.14$	2.45
	$\nu_6,\nu_5,\nu_{14}$	$6.91 \pm 0.19$	7.05	$7.06 \pm 0.13$	$7.11 \pm 0.20$	7.12
	$ u_4, \nu_3, \nu_{13}$	$0.64 \pm 0.02$	0.63	$0.61 \pm 0.03$	$0.64 \pm 0.09$	0.62
	$\nu_2,\nu_1,\nu_{12}$	$0.15 \pm 0.01$				
	$\nu_8$	$0.68 \pm 0.06$				
969 V	$\nu_7,\nu_{15}$	$2.35 \pm 0.06$				
203 K	$\nu_6,  \nu_5,  \nu_{14}$	$6.74 \pm 0.18$				

T	T - h - 1		Integra	ated band stre	$\operatorname{ength}$	
Temperature	Label	This work	Ref. $[13]^a$	Ref. $[5]^{b}$	Ref. $[6]^c$	Ref. $[7]^c$
	$ u_4,  \nu_3,  \nu_{13}$	$0.59 \pm 0.02$				
	$\nu_2,\nu_1,\nu_{12}$	$0.14 \pm 0.01$				
	$\nu_8$	$0.68 \pm 0.07$		$0.69 \pm 0.03$	$0.68 \pm 0.05$	
oro V	$\nu_7,\nu_{15}$	$2.36 \pm 0.07$		$2.41 \pm 0.08$	$2.40 \pm 0.14$	
203 K	$ u_6,  \nu_5,  \nu_{14}$	$6.76 \pm 0.19$		$6.16 \pm 0.43$	$6.96 \pm 0.20$	
	$\nu_4,\nu_3,\nu_{13}$	$0.65 \pm 0.02$		$0.65 \pm 0.05$	$0.78 \pm 0.09$	
	$\nu_2,\nu_1,\nu_{12}$	$0.13 \pm 0.02$				
	$ u_8$	$0.67 \pm 0.11$				
243 K	$\nu_7,\nu_{15}$	$2.29 \pm 0.07$				
	$\nu_6,\nu_5,\nu_{14}$	$6.58 \pm 0.18$				
	$\nu_4,\nu_3,\nu_{13}$	$0.64 \pm 0.02$				
	$\nu_2,\nu_1,\nu_{12}$	$0.13 \pm 0.02$				
022 IZ	$ u_8$	$0.67 \pm 0.15$		$0.71 \pm 0.01$		0.75
	$\nu_7,\nu_{15}$	$2.35 \pm 0.07$		$2.30 \pm 0.04$		2.51
200 K	$ u_6,  \nu_5,  \nu_{14}$	$6.92 \pm 0.21$		$6.19 \pm 0.02$		7.38
	$\nu_4,\nu_3,\nu_{13}$	$0.61 \pm 0.02$		$0.62 \pm 0.05$		0.64
	$\nu_2,\nu_1,\nu_{12}$	$0.13 \pm 0.01$				
	$ u_8$	$0.64 \pm 0.15$				
000 IZ	$\nu_7,\nu_{15}$	$2.30 \pm 0.08$				
220 K	$\nu_6,  \nu_5,  \nu_{14}$	$6.67 \pm 0.19$				
	$\nu_4,\nu_3,\nu_{13}$	$0.58 \pm 0.02$				
	$\nu_2,\nu_1,\nu_{12}$	$0.12 \pm 0.01$				

Table 2 – continued from previous page

Temperature	Labol		Integra	ted band stre	ength	
Temperature	Laber	This work	Ref. $[13]^a$	Ref. $[5]^{b}$	Ref. $[6]^c$	Ref. $[7]^c$
	$ u_8$			$0.70 \pm 0.07$		
213 K	$\nu_7,\nu_{15}$			$2.46 \pm 0.05$		
	$\nu_6,  \nu_5,  \nu_{14}$			$6.80 \pm 0.28$		
	$\nu_4,\nu_3,\nu_{13}$			$0.60 {\pm} 0.04$		
	$\nu_8$			$0.79 \pm 0.09$		
909 IZ	$\nu_7,\nu_{15}$			$2.34 \pm 0.26$		
203 K	$\nu_6,\nu_5,\nu_{14}$			$6.90 \pm 0.10$		
	$\nu_4,\nu_3,\nu_{13}$			$0.58 {\pm} 0.03$		

Table 2 – continued	from	previous	$\mathbf{page}$
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